

WASP the wide angle neutron spin echo instrument

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ESS-ILL 2022





Lund 7 Oct 2022

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+50 others !!

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Acknowledgments



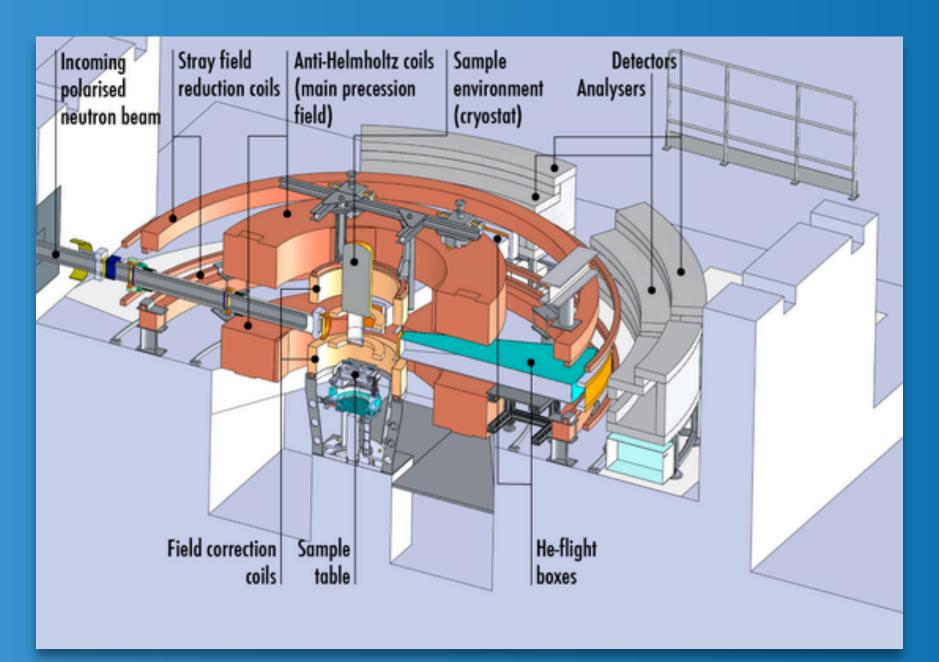




HZ Berlin Beate Brüning Robby Krischnik Stefan Wellert



TU Delft Katia Pappas







Acknowledgments



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Why am here ?

Everything was said the first day:

we eliminate scanning ? (Martin)

 WASP is ILL's flagship Spin Echo we have the first results (Jacques, Helmut)

ESS has no NSE 'capability gap' (Andreas)









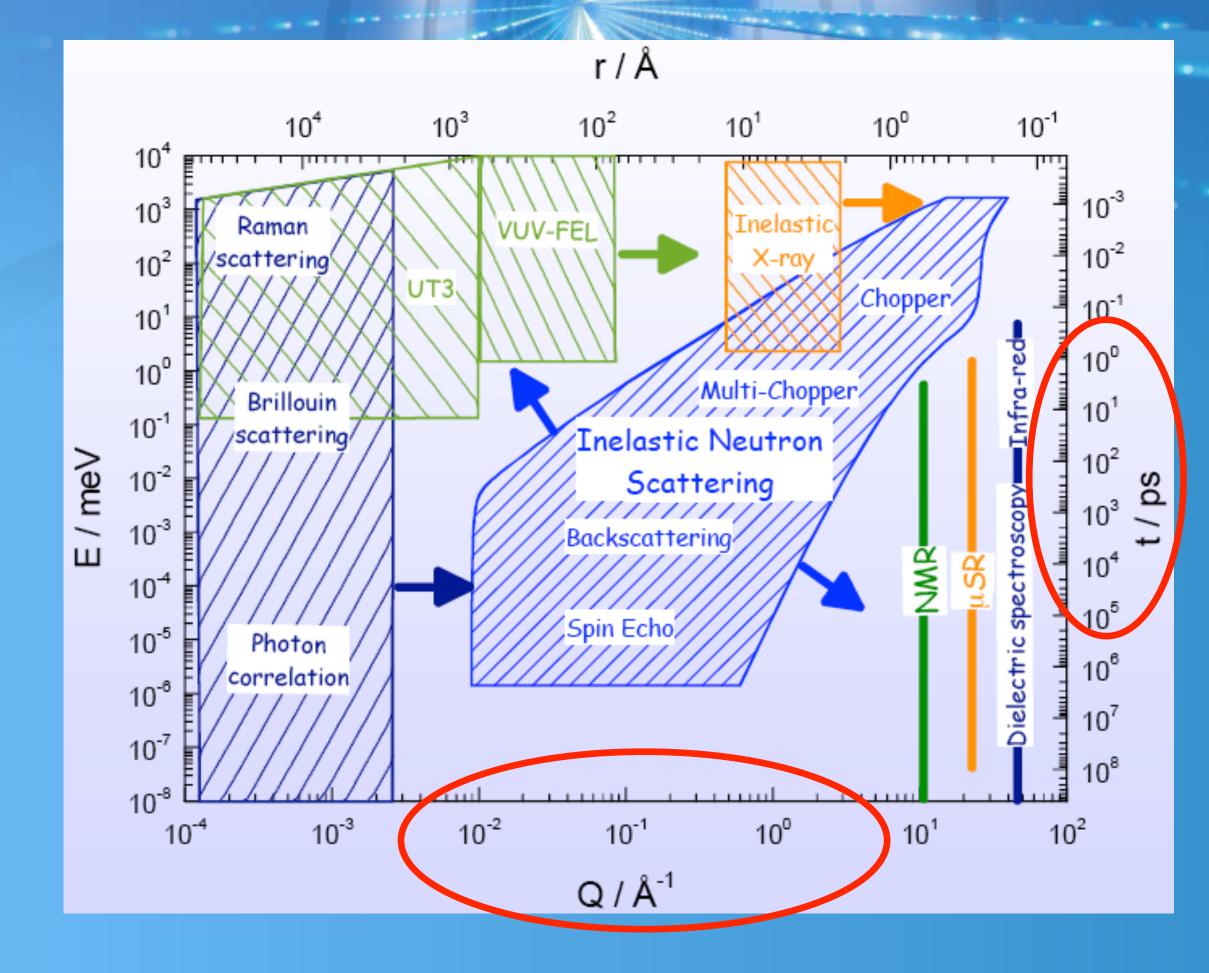
- High resolution techniques with scanning are slow, can







Why NSE important



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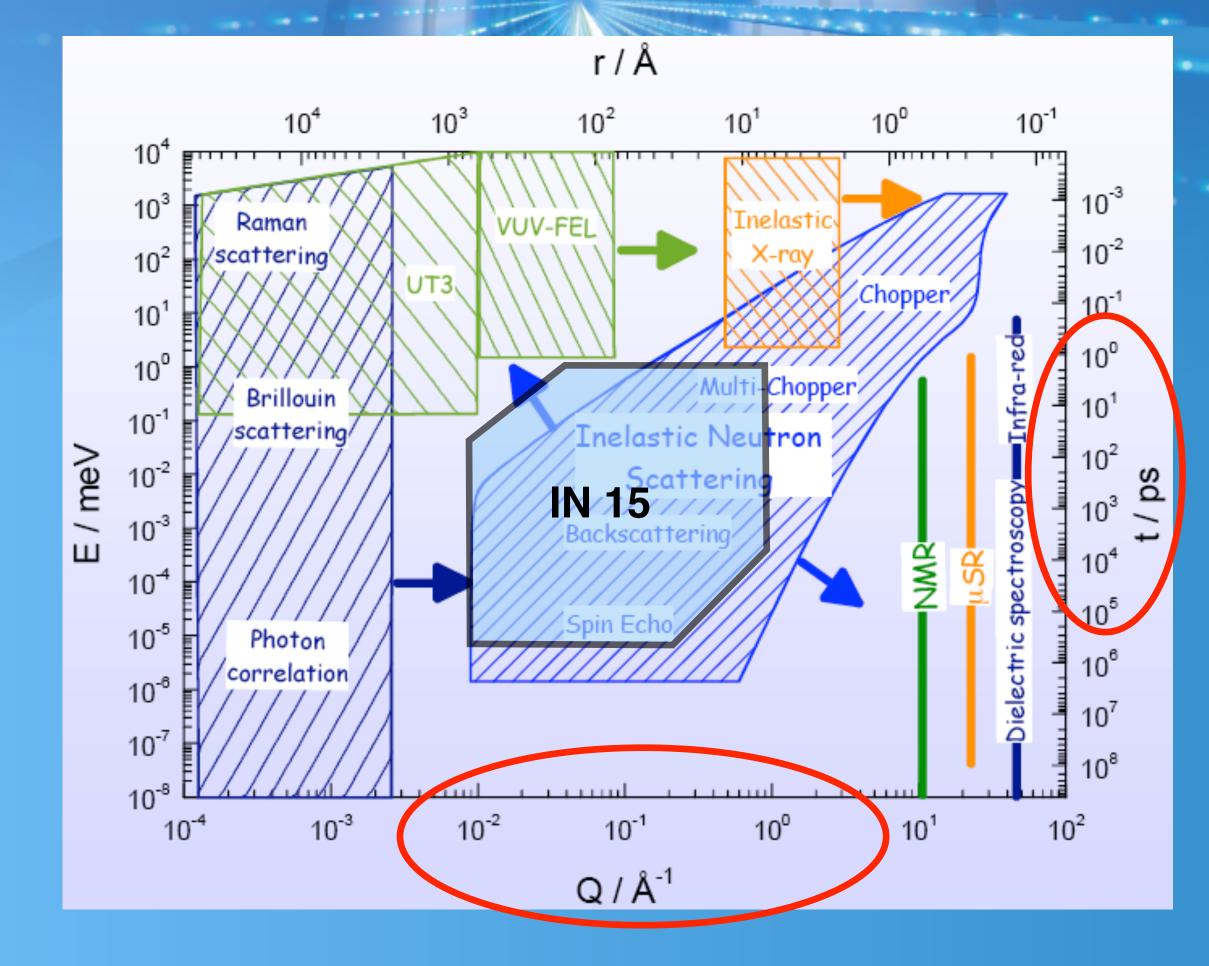


- Unmatched Q-t range •
- Works in time space not energy space
- Sees difference not sum of coherent and incoherent scattering
- For magnetism XYZ polarisation analysis built in





Why NSE important



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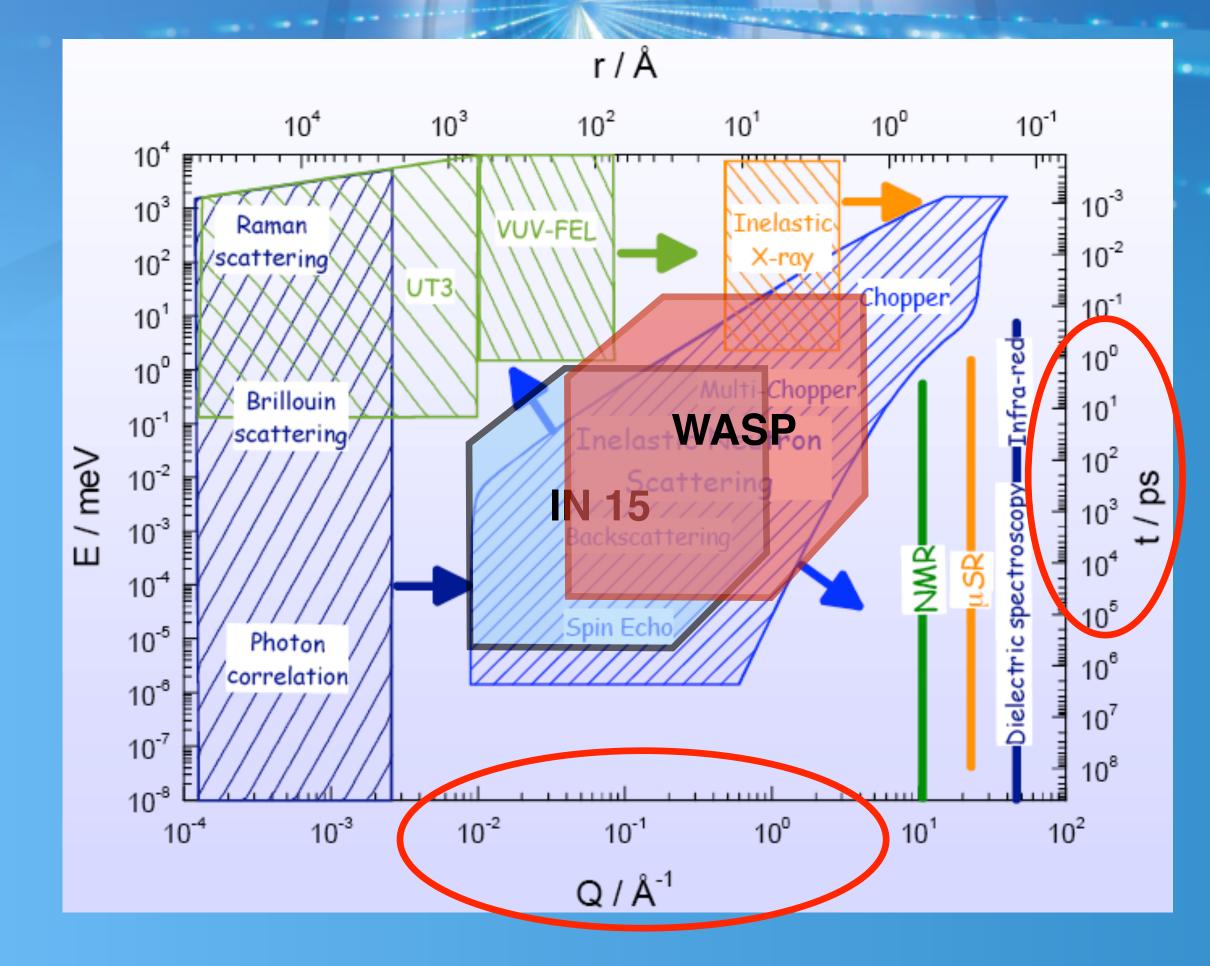
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• IN 15 is optimized for molecular length scales 0.01-1 Å⁻¹, 5ps-1000 ns





Why NSE important



• WASP is optimized for atomic to molecular length scales 0.1-4 Å⁻¹, 0.2 ps-100 ns

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indamentals of NSE

We want to measure the difference The classical method is defining the final and initial velocity Defining 2x = throwing away all other neutrons 2xHigh resolution == very few neutrons remain Can we use all neutrons without defining/monochromatizing?

Yes ! We will use the neutron spin (Feri Mezei 1972)

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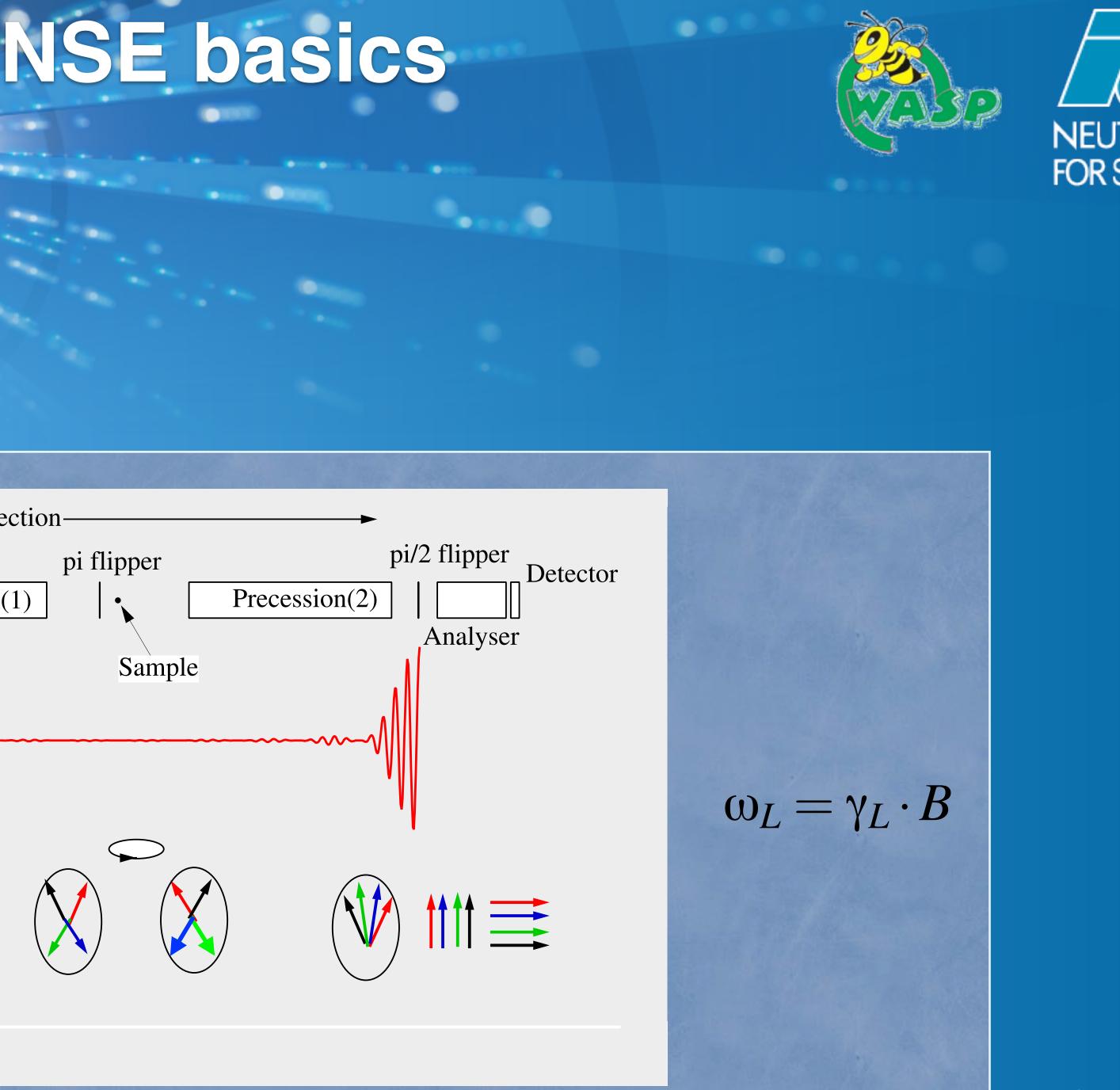
 $\hbar ω = m v^2/2 - m v^2/2$

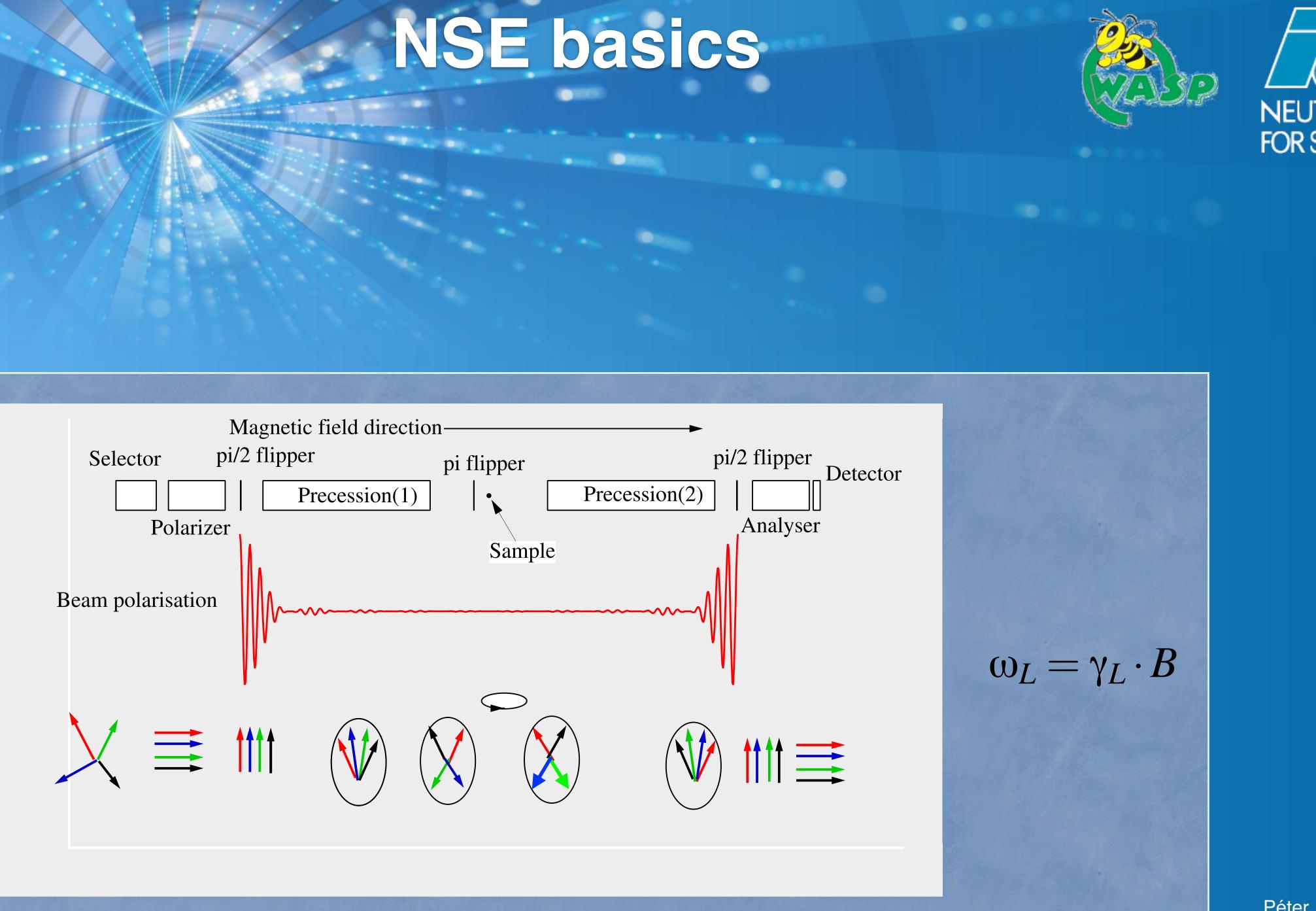
final velocity

initial velocity



















For elastic scattering:

For omega energy exchange:

 $\varphi_{tot} = \frac{\gamma B}{T}$ $\varphi_{tot} = \frac{\hbar \gamma B l}{m v^3} \omega$

 $\langle \cos \varphi \rangle = -$

S(

The probability of omega energy exchange:

The final polarization:









Echo condition:

$$\int_{\pi/2}^{\pi} B_1 d\ell = \int_{\pi}^{\pi/2} B_2 d\ell$$

For elastic scattering:

For omega energy exchange:

The probability of omega energy exchange:

The final polarization: $\langle \cos \varphi \rangle = \frac{J}{-1}$

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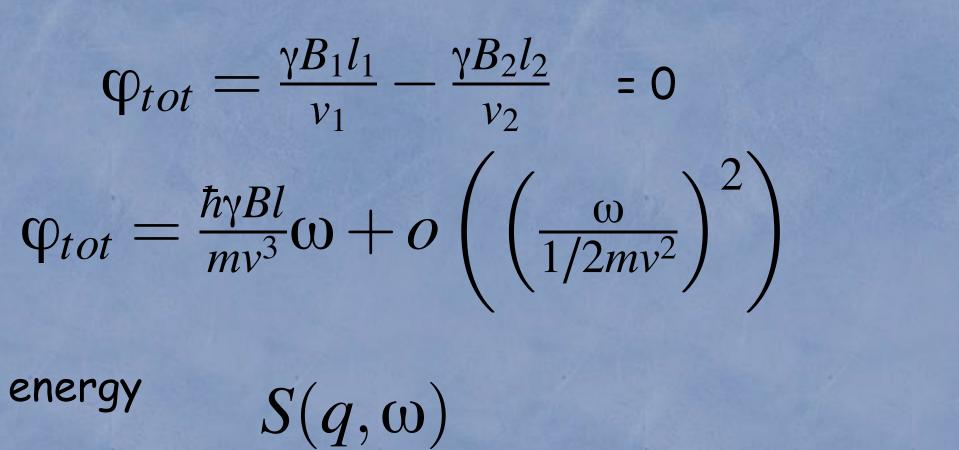
NSE basics





The measured quantity is: S(q,t)/S(q,0)where

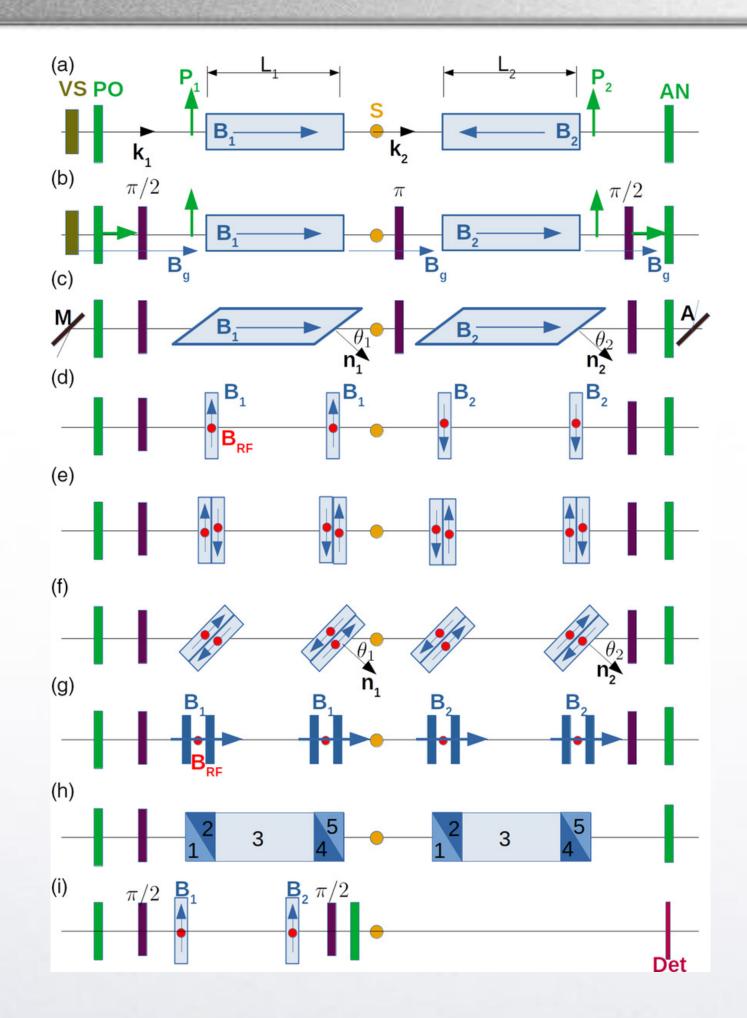
$$t \propto \lambda^3 \int B d\ell$$



$$\frac{\cos(\frac{\hbar\gamma Bl}{m\nu^3}\omega)S(q,\omega)d\omega}{\int S(q,\omega)d\omega} = S(q,t)$$



Other Echo methods



Keller T et al. Physica Status Solidi (b), Volume: 259, Issue: 5, (2021), DOI: (10.1002/pssb.202100164)

We can code energy or angle or any combination of the two.

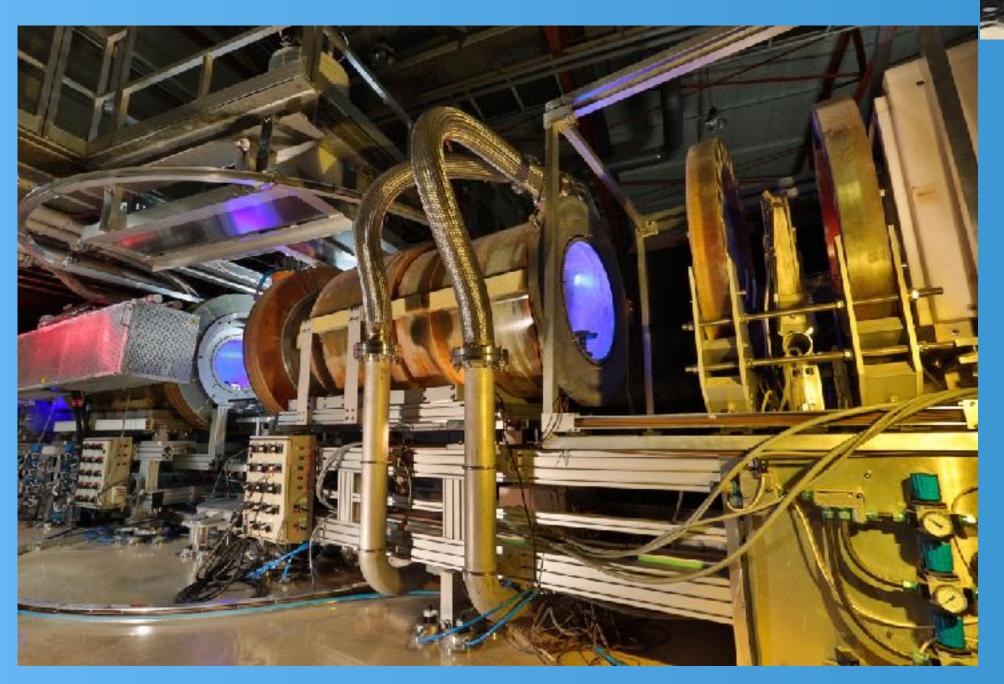
Either magnetic filed is static and neutron rotates or neutron is static in 0 field and magnetic field rotates





NSE Spectrometers at the ILL

IN15



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IN11 (1972-2020)

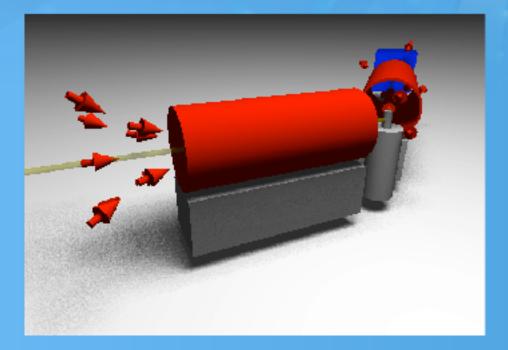




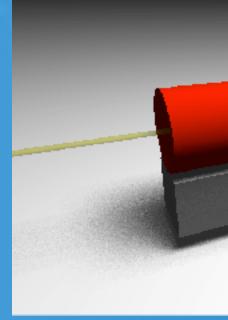


IN11A - high res Mezei '77, Farago '92

IN11C - 30° detector bank Farago '97







WASP will:

- Provide same high resolution as IN11A/ old IN15
- Increase the sample flux x8
- Increase the detection solid angle x3 (90° compared to IN11C)

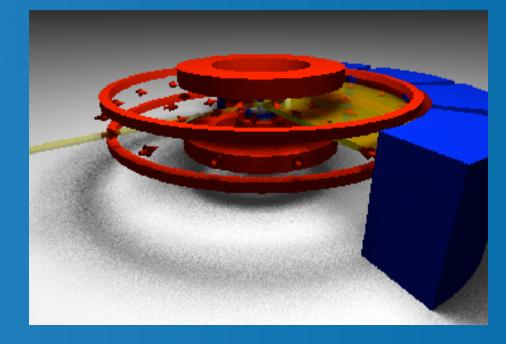
NSE Evolution





SPAN 30° detector bank Mezei Pappas '99

WASP 90° detectors '18



25x higher intensity and 6x higher field integral than IN11C

x**25**

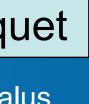
Full q range results on hydrogenated samples in <u>hours</u> not weeks

Images Courtesy of P. Fouquet

Reference: J. Neutron Res. 15, 39 (2007)

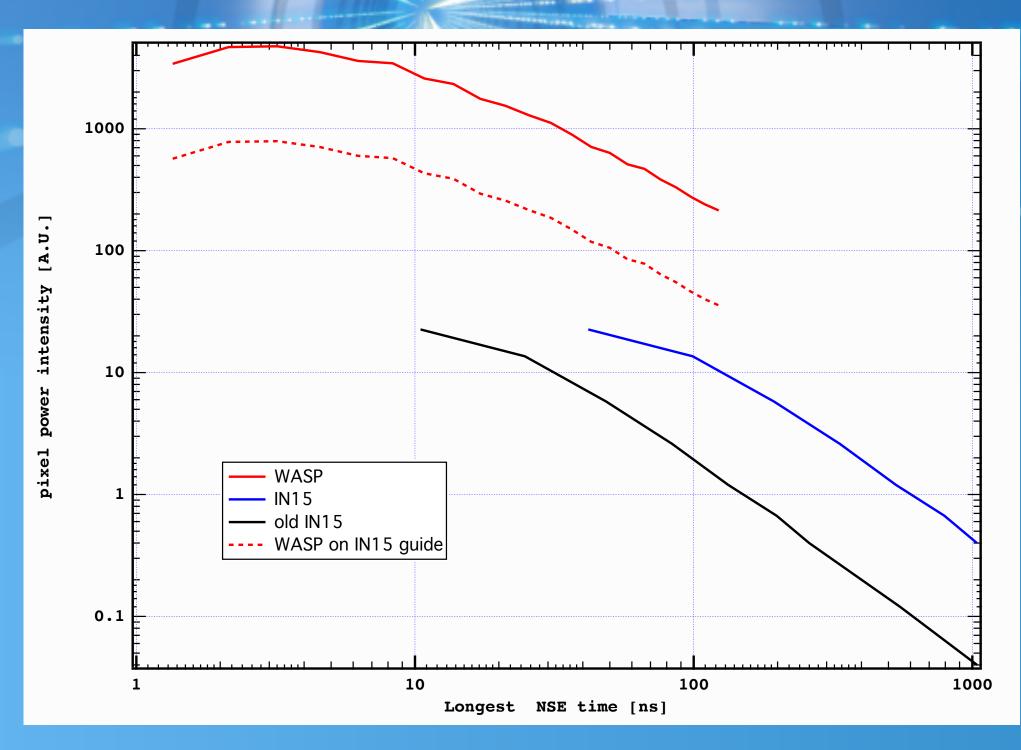
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- Wide Q-range : Sample environments are limited - No ferromagnetic/intensity modulated echo - Bigger footprint - Construction cost (polarisers)

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WASP advantages





✓ High <u>pixel</u> intensity where overlaps with small angle NSE

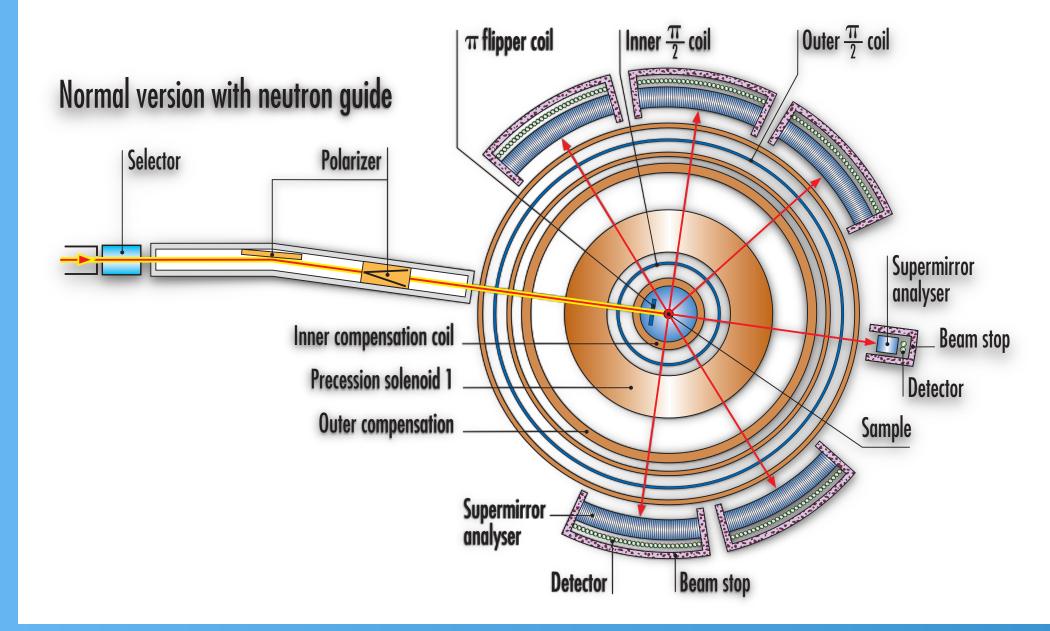
✓ Wide Q-range less tuning, less scanning less overhead \checkmark Kinetic studies at a pulsed source ✓ Usable as polarised diffractometer (D7)





Wasp magnetic field layout (SPAN)



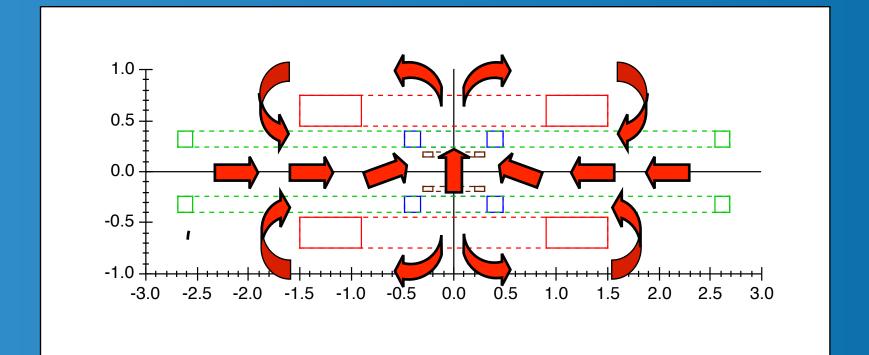


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5 years mirror production 240m² coating • 90 casettes 3300 mirrors for 90 degrees • Permanent magnet produced field



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Analysers (D7)













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120

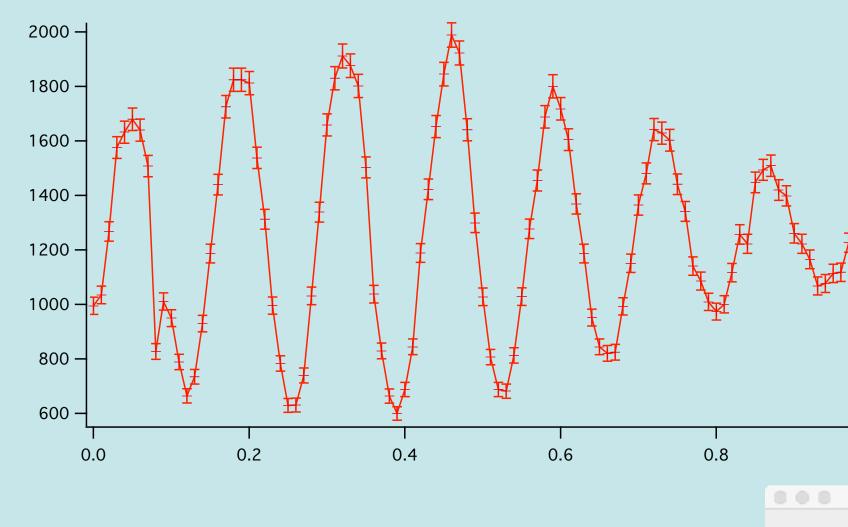
100

80 -

60 -

40 -

20 -



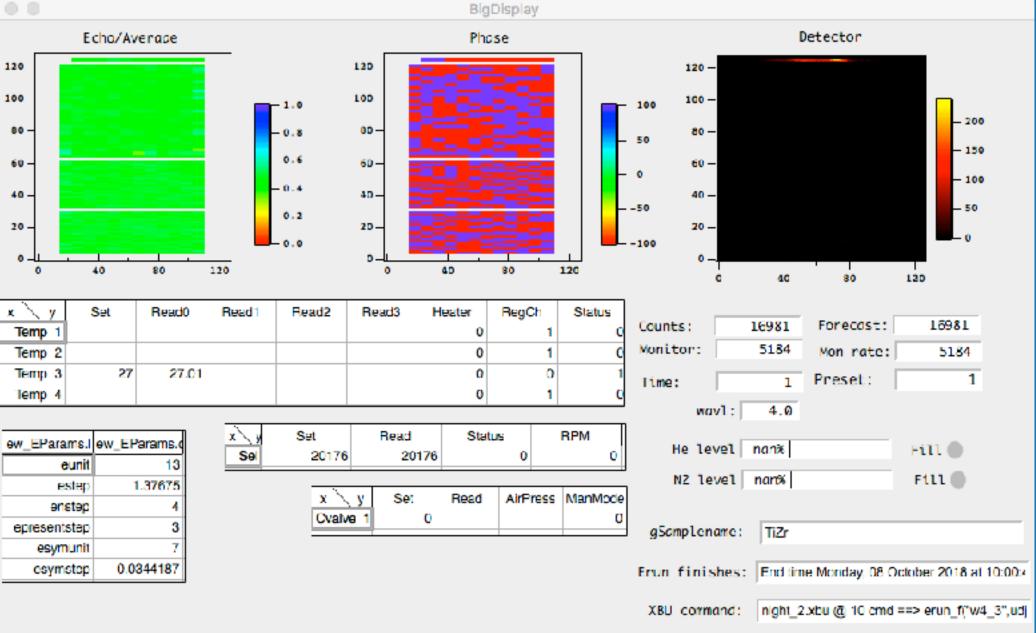
Full intensity authorised, echo in all detectors 10/2018

First Results





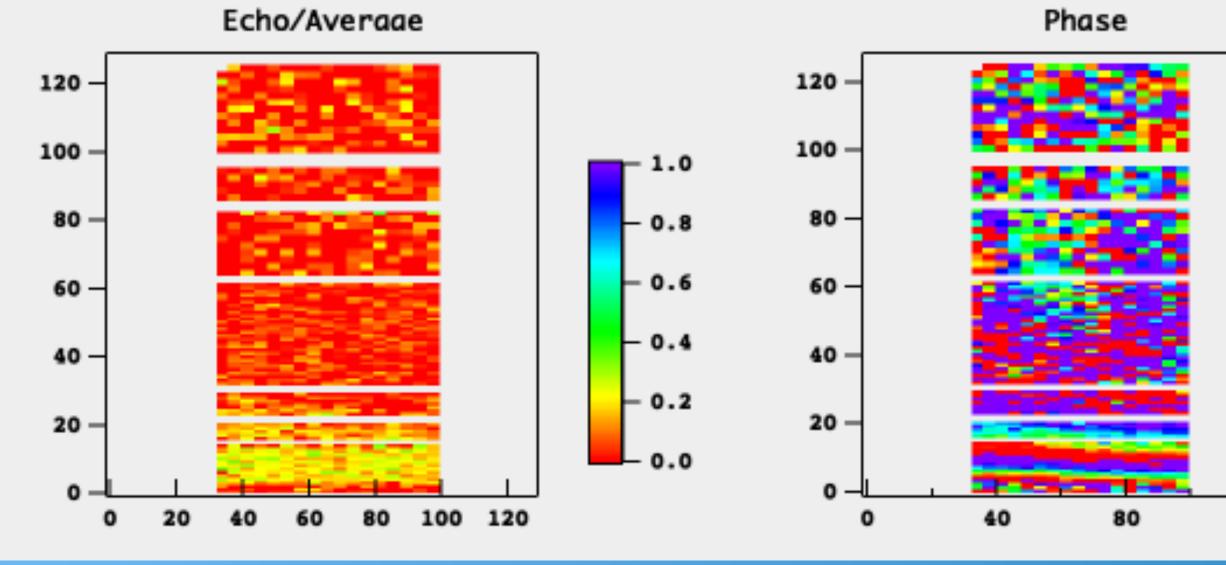
Security file signed 3/10/2018 (limited to 1/5 intensity) First echo 4/10/2018



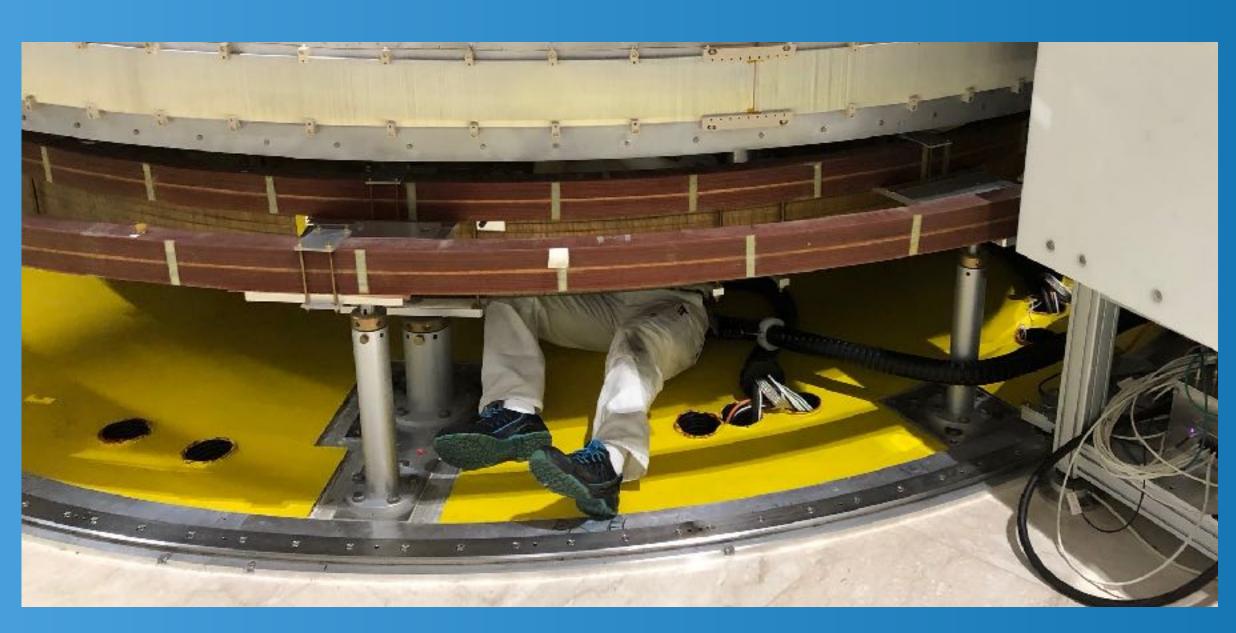








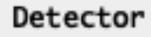
3x Mechanical Magnetic Neutrons

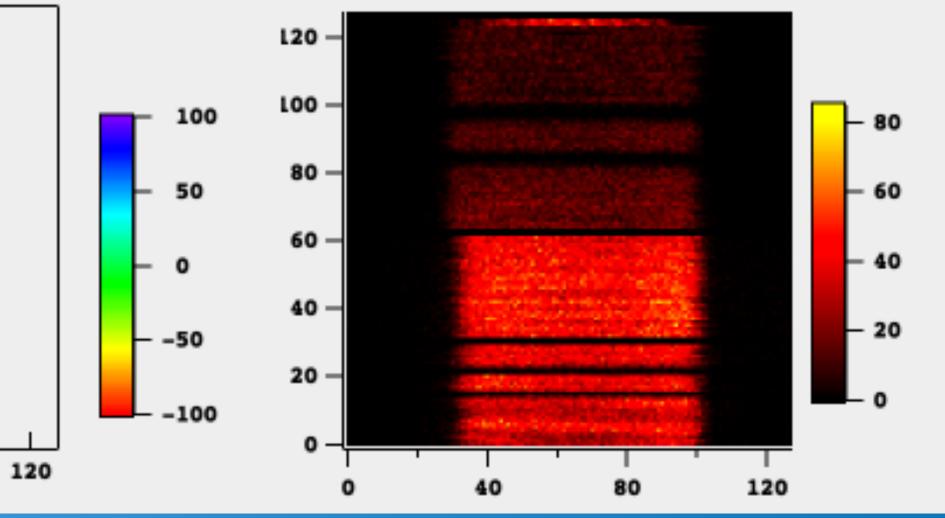


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Tuning







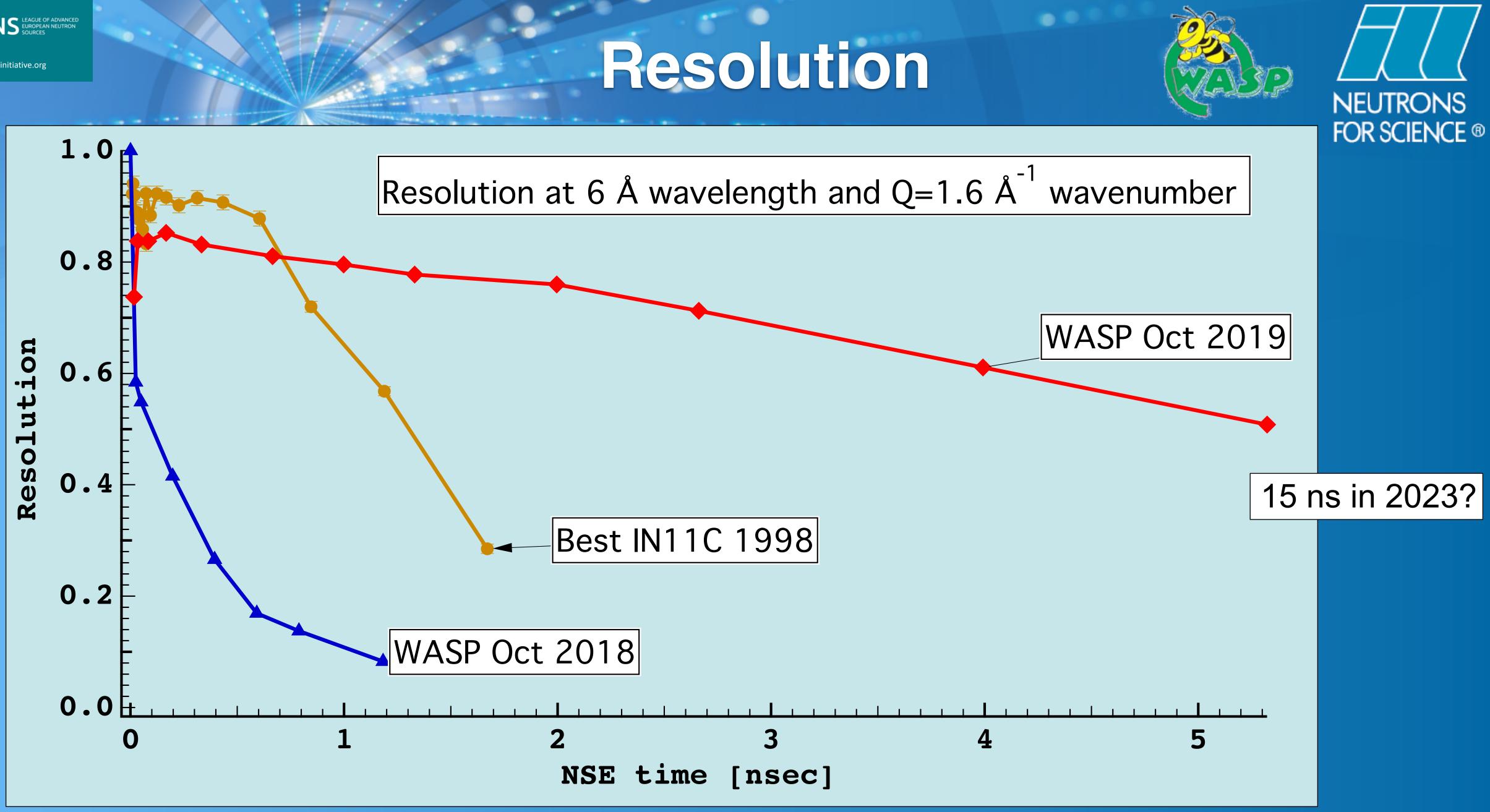
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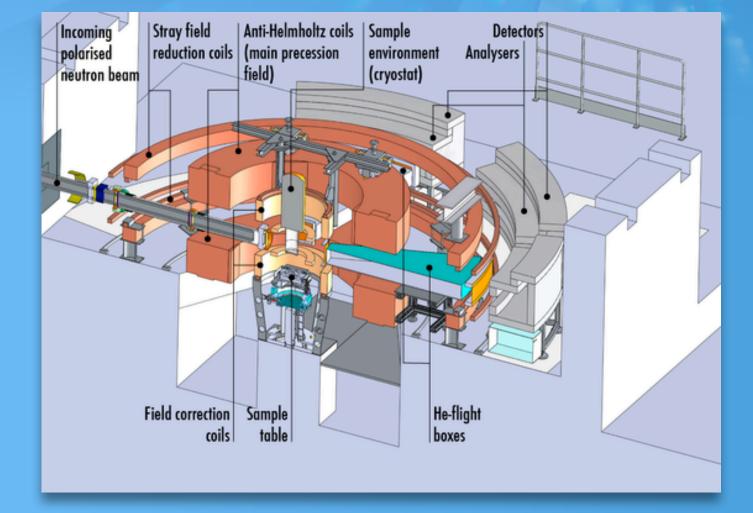
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WASP specs





•dynamic range 0.2 ps - 100-ns •0.05 - 4 Å⁻¹ •3-14 Å wavelength •signal = 500x IN11A •~ 50 t Cu •~ 0.6 MW max power, 50 kW average power

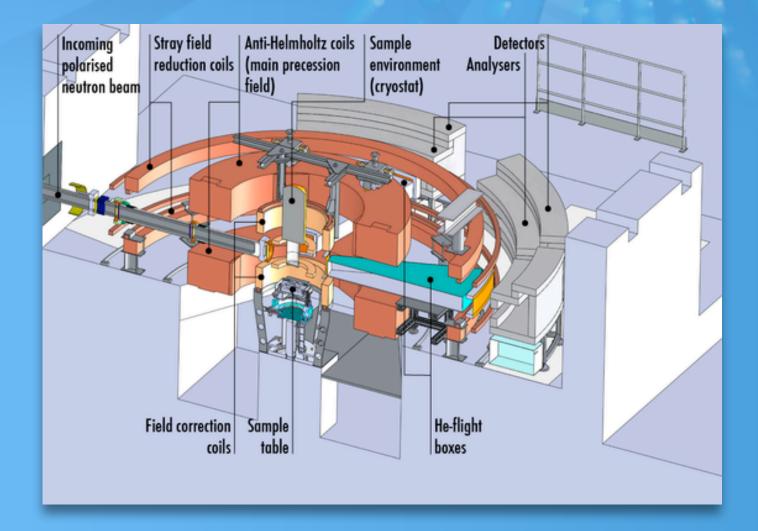








WASP specs



First draft of proposal 2001 Proposal to instrument subcommittee 2005 Coil manufacturing 2015 Polarisation tests Apr 2018 First echo 4 October 2018

First users end of 2019 2020

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 dynamic range 0.2 ps - 100 ns •0.05 - 4 Å⁻¹ •3-14 Å wavelength •signal = 500x IN11A •~ 50 t Cu •~ 0.6 MW max power, 50 kW average power



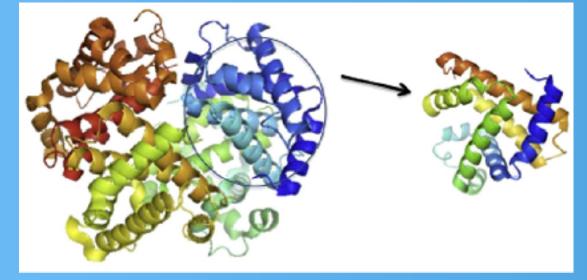






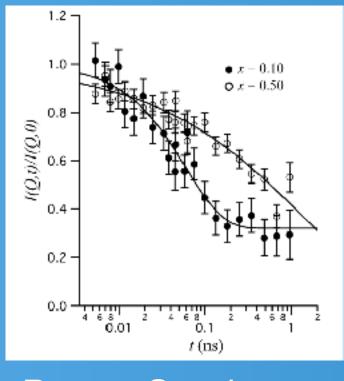
Scientific Case for WASP

Functionality of Biomolecules



Hemoglobin Activity (incoherent) J. Lal et al., J. Mol. Biol. **397**, 423 (2010)

Energy Materials Incoherent Scattering



Proton Conductors M. Karlsson et al., J. Phys. Chem. C **114**, 3292 (2010)

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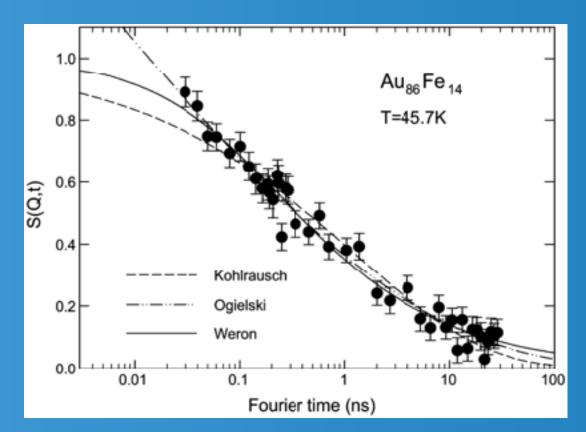
Courtesy of P Fouquet





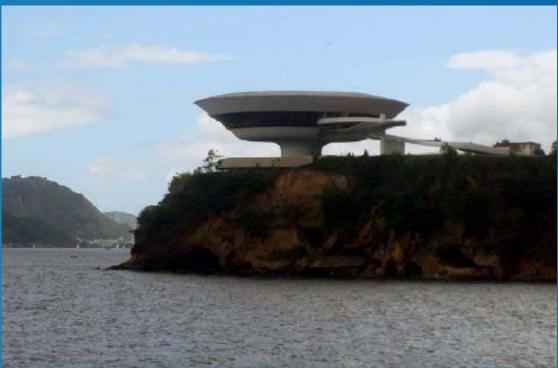
NOT for coherent simple diffusion

Glass Transition



Relaxation Function Test on Spin Glasses R.M. Pickup et al., PRL 102, 097202 (2009)

Liquids at the Nanometer Scale, Confinements



Water mobility in concrete

H. Bordallo et al., ACS Applied Materials & Interfaces 1, 2154 (2009)

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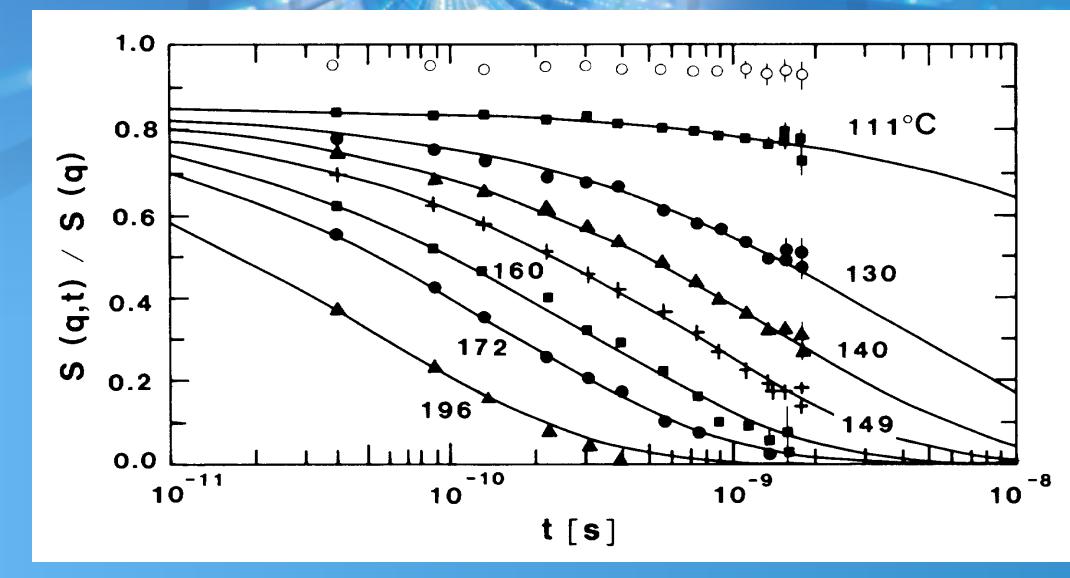








CKN comparison



Neutron spin echo study of dynamic correlations near the liquid-glass transition

F. Mezei, W. Knaak, and B. Farago Phys. Rev. Lett. 58, 571 – IN11. 1987.

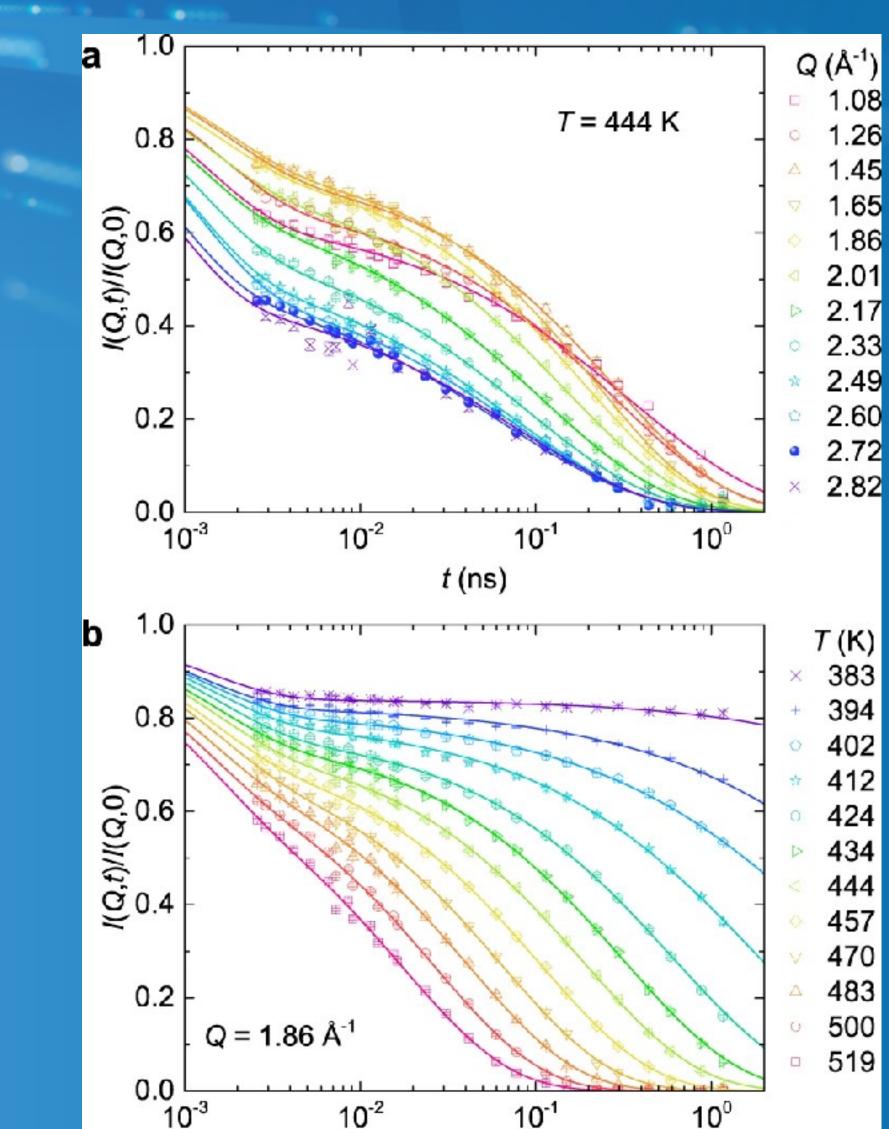
Q-dependent collective relaxation dynamics of glass-forming liquid Ca0.4K0.6(NO3)1.4 investigated by wide-angle neutron spin-echo

P. Luo, Y. Zhai, P. Falus, V. García Sakai, M. Hartl, M. Kofu, K. Nakajima, A. Faraone, YZ. Nature Communications, 13, 2092 (2022). DOI: 10.1038/s41467-022-29778-4 WASP 2022

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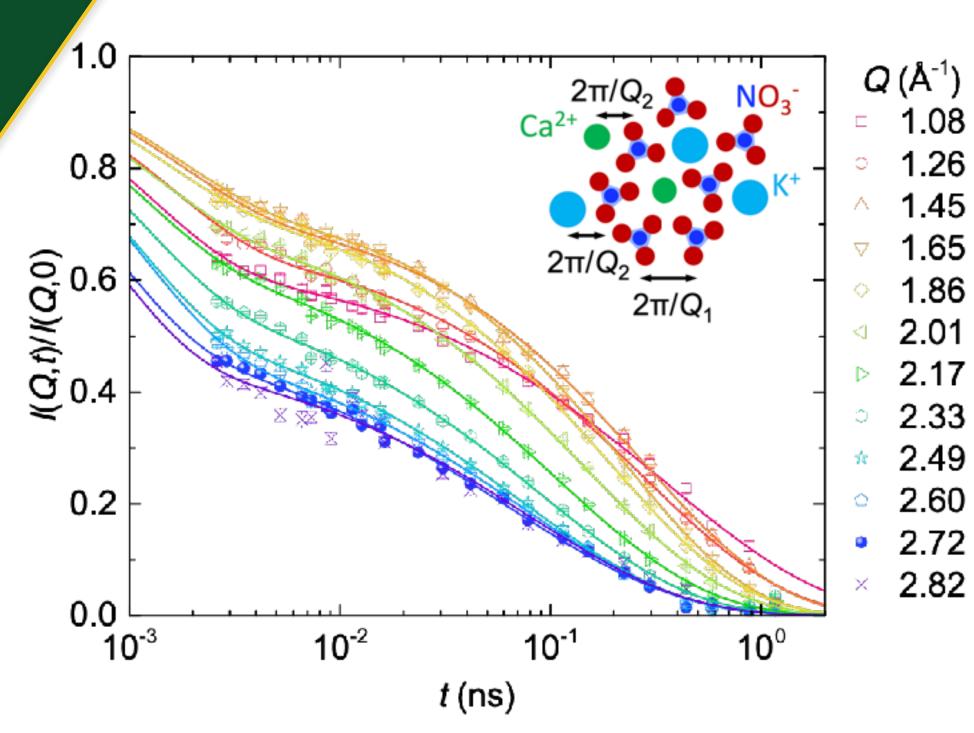




t (ns)



New insight into the collective relaxation dynamics of glassforming liquid revealed by wide-angle neutron spin echo



Wide-angle neutron spin echo measurement of a model fragile liquid $Ca_{0.4}K_{0.6}(NO_3)_{1.4}$ (CKN) at T=444 K at various wavevectors (Q). The inset is an illustration of CKN structure.

P. Luo, Y. Zhai, P. Falus, V. García Sakai, M. Hartl, M. Kofu, K. Nakajima, A. Faraone, YZ. Nature Communications, 13, 2092 (2022). DOI: 10.1038/s41467-022-29778-4

Neutron facilities at the ILL, NCNR, and J-PARC were used.



Scientific Achievement

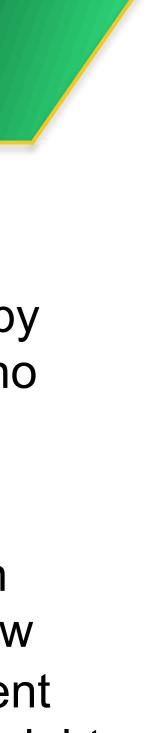
The collective relaxation dynamics over the full range of the microscopic structural length scales have been characterized by employing the recently developed wide-angle neutron spin-echo (WASP) spectroscopy.

Significance and Impact A change in the dominant relaxation mechanisms was found in liquid $Ca_{0.4}K_{0.6}(NO_3)_{1.4}$ (CKN) at the length scale of 2.6 Å, below which the relaxation process exhibits a temperature independent distribution and more Arrhenius-like behavior, revealing new insight into the collective dynamics in glass-forming liquids – a key to understand the universality of liquid state physics.

Research Details



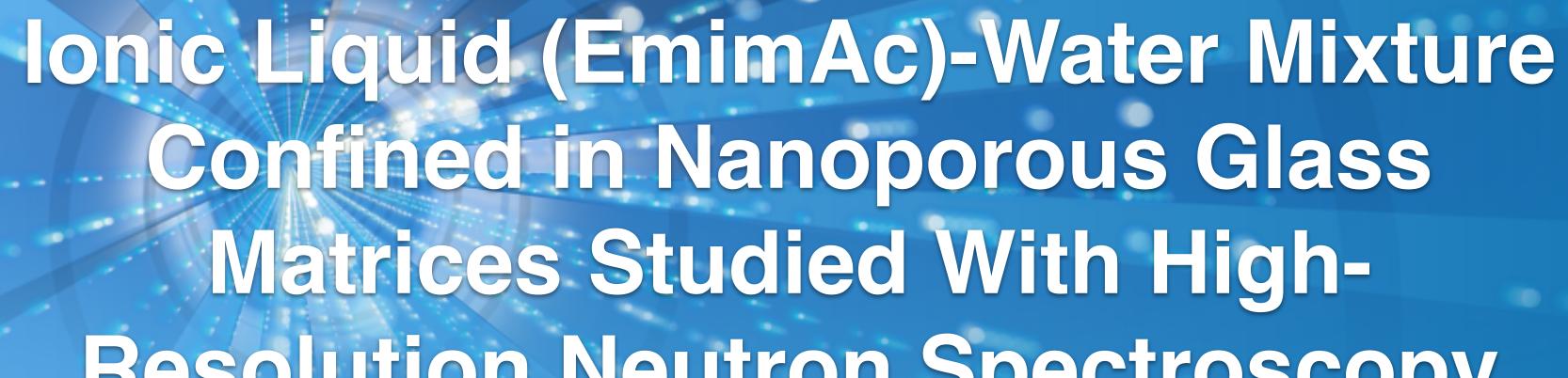
CKN sample was prepared by mixing high-purity $Ca(NO_3)_2 \cdot 4H_2O$ and KNO_3 and then drying the mixture in vacuum above the melting point. Intermediate scattering functions of liquid CKN at various temperatures from 383 K to 519 K, and in the Q-range from 1.08 $Å^{-1}$ to 2.82 $Å^{-1}$ were measured on WASP at the ILL.



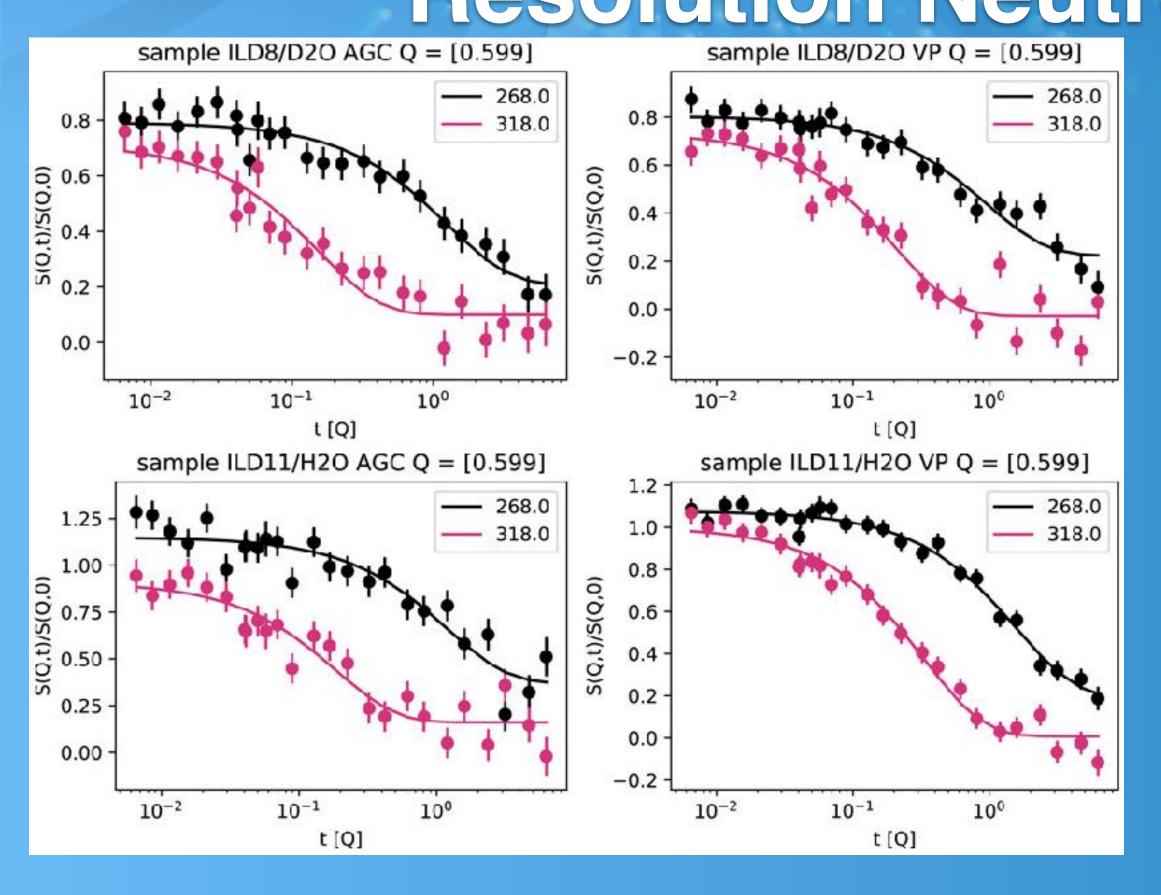












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Resolution Neutron Spectroscopy H. Frielinghaus, M. Fomina, D. Hayward, P. S. Dubey, S. Jaksch, P. Falus, P. Fouquet, L. Fruhner and O. Holderer

Front. Phys., 21 April 2022 Sec. Physical Chemistry and Chemical Physics https://doi.org/10.3389/fphy.2022.872616

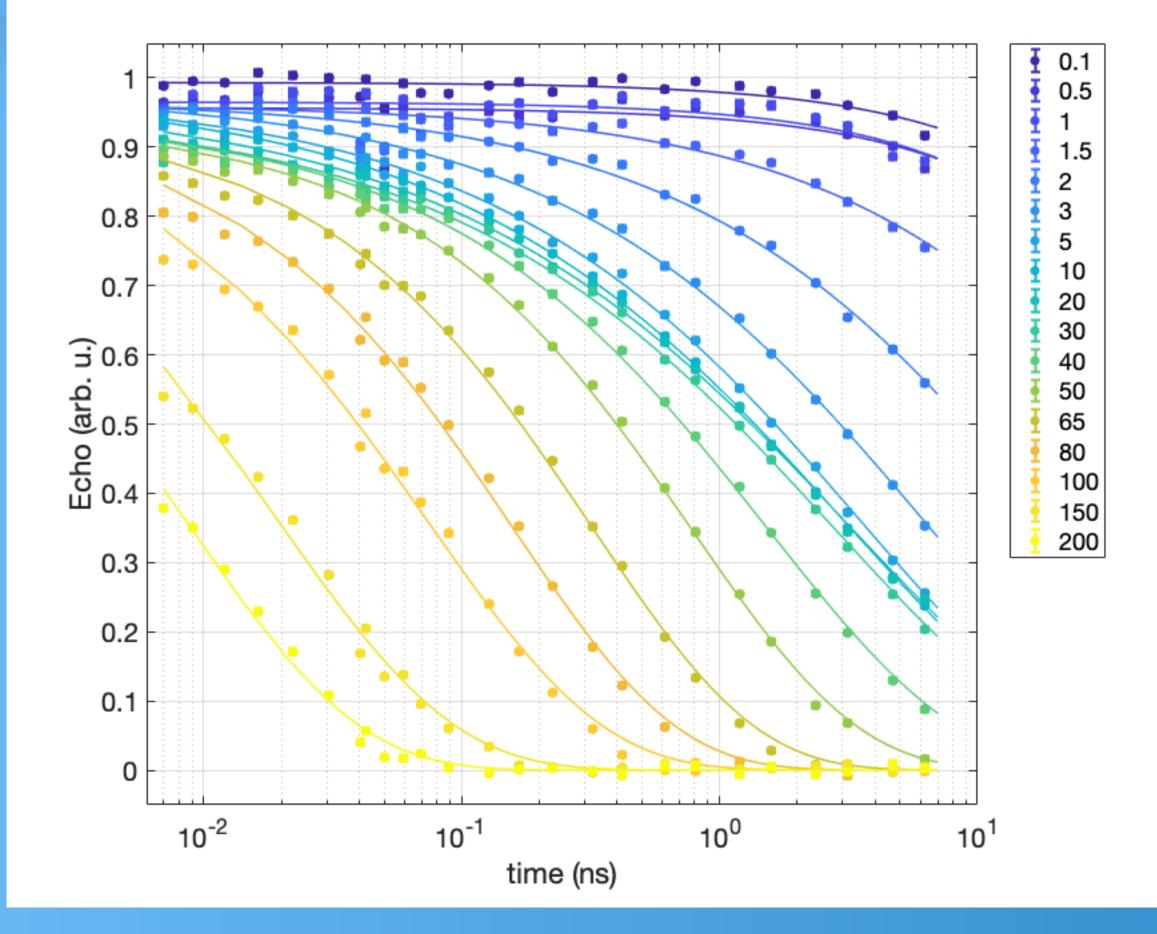






More examples

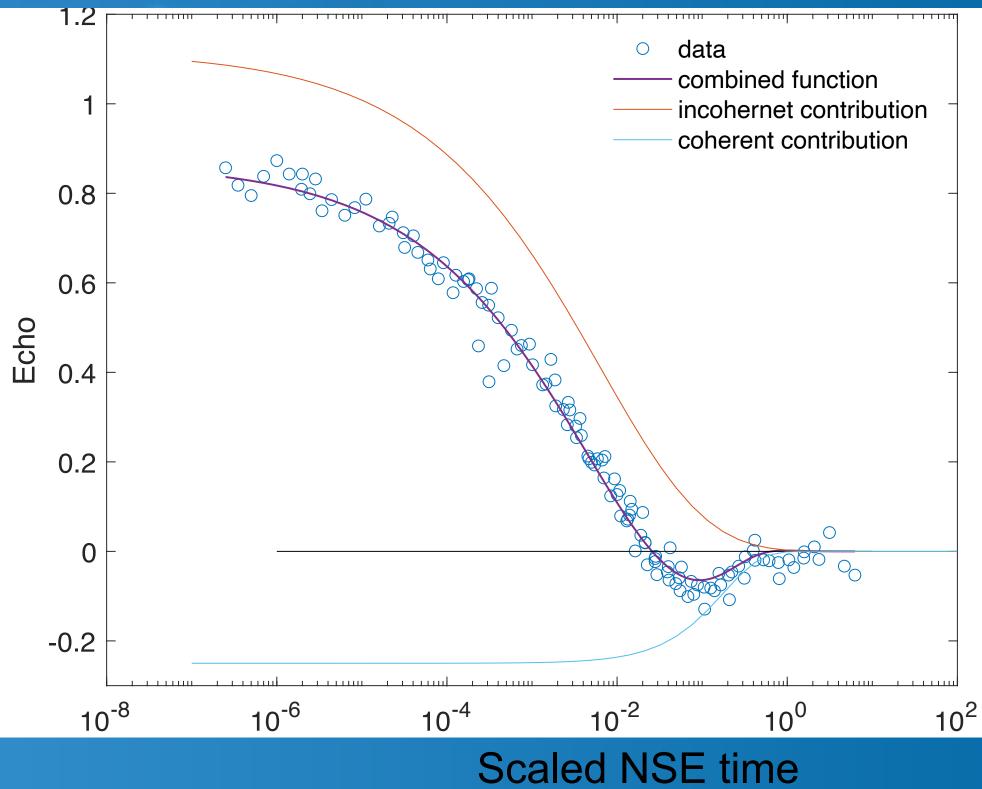
Doped HoTiO spin glass



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Diluted ionic liquid electrolyte



J. Phys. Chem. C (2022), 126, 38, 16262–16271













Conclusion/ Homework

 Wasp sees things we have never seen before, apply for ILL beamtime ! (Result in hours not a week)

 Need higher resolution than time of flight or backscattering? Need more intensity at high Qs than small angle NSE ? Apply for beam time !

ESS has no Neutron Spin Echo, lobby for one !



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