

# OPPORTUNITIES IN THE USE OF VERY COLD NEUTRONS FOR REFLECTOMETRY TECHNIQUES.

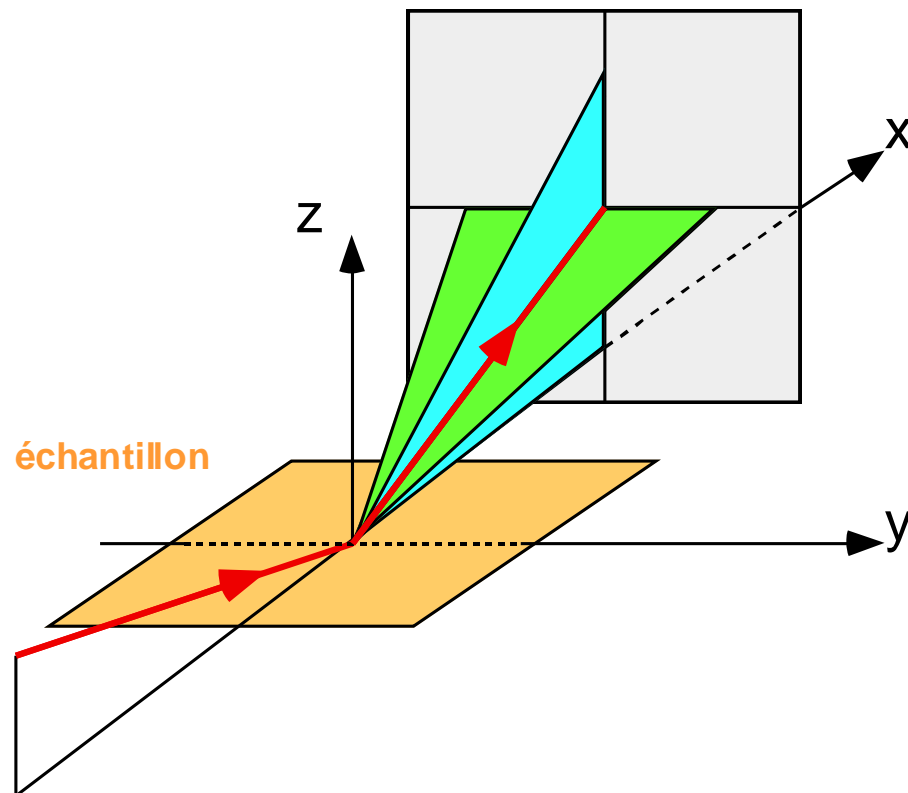
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# WHAT DO WE AIM AT PROBING ?

## What science ?

- → nanoscale (1-10nm) structure of interfaces
  - Lower values : not accessible
  - Higher values : of little interest
- Only a structural method (so far)
- Polymer films interfaces, solid liquid interfaces, liquid surfaces, magnetic films



- **Specular reflection**  
 $0.06 < Q_z < 3 \text{ nm}^{-1}$   
 $1 \text{ nm} < \xi < 100 \text{ nm}$
- **Incidence plane (Off-specular)**  
 $10^{-4} < Q_x < 10^{-2} \text{ nm}^{-1}$   
 $600 \text{ nm} < \xi < 60 \text{ }\mu\text{m}$
- **⊥ incidence plane (GISANS)**  
 $10^{-4} < Q_y < 3 \text{ nm}^{-1}$   
 $3 \text{ nm} < \xi < 100 \text{ nm}$

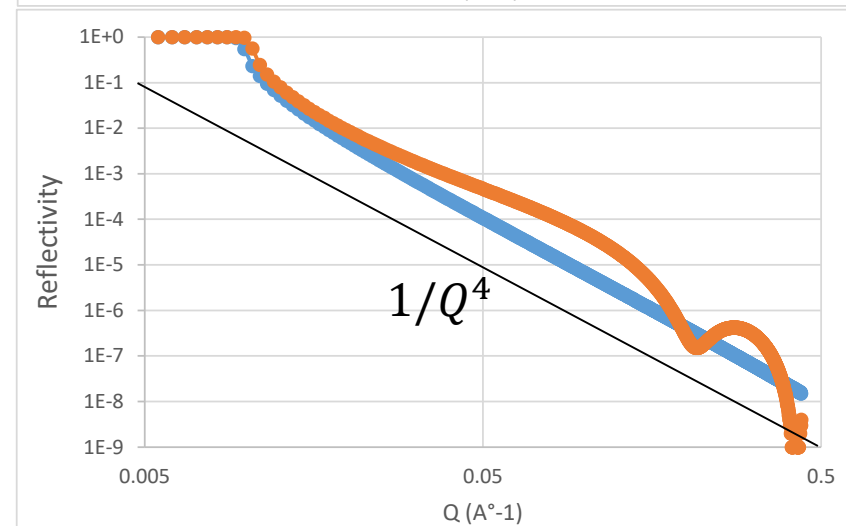
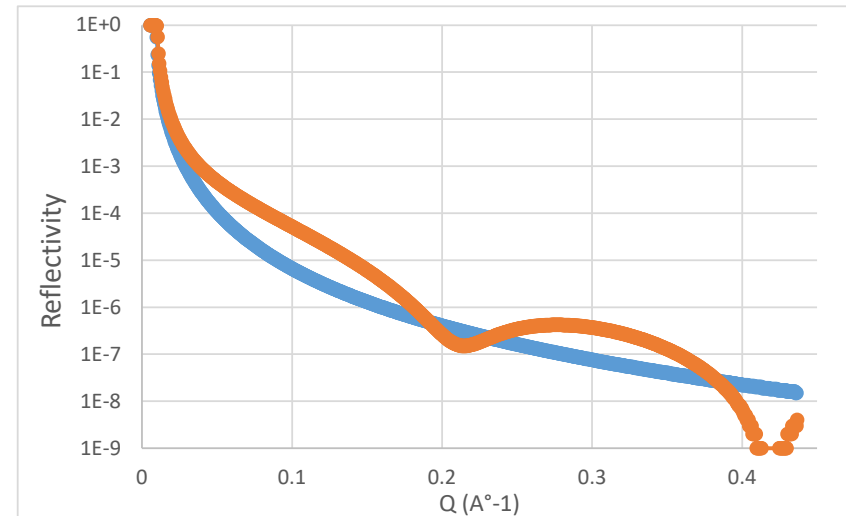
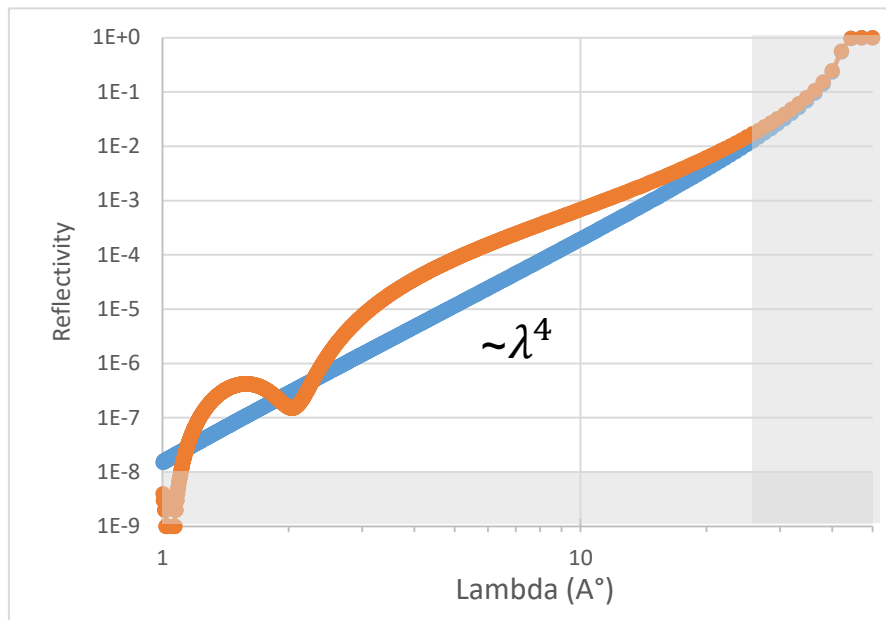
# REFLECTOMETRY IS ESPECIALLY SUITED TO TOF

NOTA: the presentation focusses on VNC use at pulsed sources

- Sample reflectivities follow a  $1/Q^4$  trend with modulations around this general trend

$$R(Q) \propto 1/Q^4$$

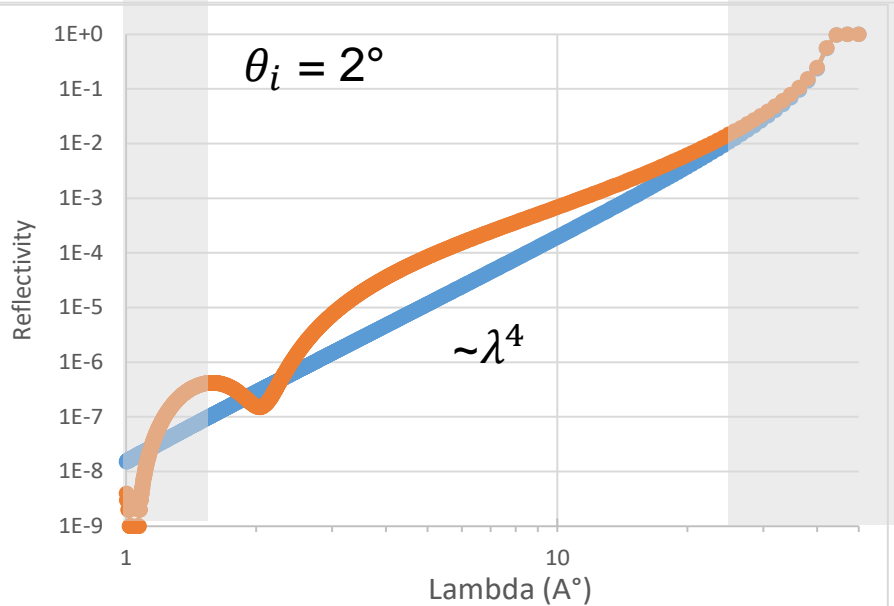
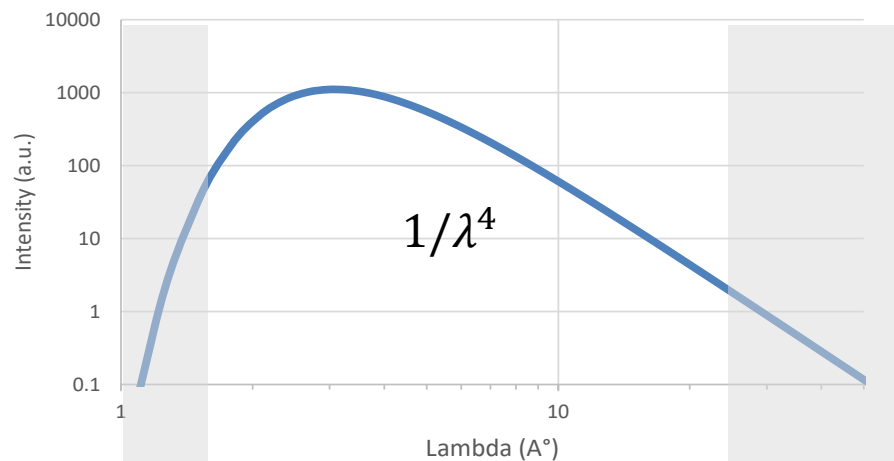
- In « experimental space » the measurement looks as (incidence angle of  $2^\circ$ )



Si substrate and  $\text{SiO}_2/\text{Si}$

# MATCHING THE INCIDENT SPECTRUM AND THE REFLECTIVITY RANGE OF INTEREST

Maxwell Boltzmann (@50K)



Reflectometry measurements on a long pulse source are using neutrons very efficiently in TOF mode.

- All neutrons provide information with a roughly equivalent statistics over a wide Q range (within a factor 3)

$$Q = \frac{4\pi}{\lambda} \sin \theta_i$$

The probed Q-range of interest can be tuned by simply changing the incidence angle  $\theta_i$

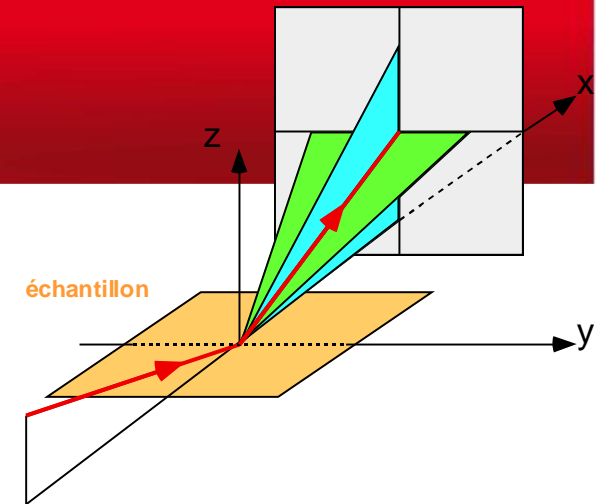
# WHAT IMPACT OF USING VERY COLD NEUTRONS ?

- 1°/ The physics probed by the neutrons does not depend on the wavelength
- 2°/ For a given Q range of interest
$$Q = \frac{4\pi}{\lambda} \sin \theta_i \approx \frac{4\pi}{\lambda} \theta_i$$
  - If the wavelength spectrum is shifted by a factor  $\alpha$ , it involves using incidence angles  $\theta_i$  which are  $\alpha$  times larger.
  - Assuming a constant  $\delta\theta/\theta_i$ , the angular resolution can be relaxed by a factor  $\alpha$  (by simply relaxing the collimation slits)
  - There are second order corrections if the sample is over-illuminated or under-illuminated
- ➔ The gain in flux is proportionnal to the wavelength spectrum shift  $\alpha$
- Note that for large samples (ex. Langmuir trough 100mm long), increasing the incidence angle from typically 4° to 12°, requires designing collimations which are on the order of 12m to fully benefit from the new wavelength spectrum ➔ It is not a technical challenge but it is an unusual design for reflectometers.
- 3°/ The absorption effects remain usually negligible
- 4°/ There are no « multiple scattering » issues. The measurement remains as « clean and simple » as before whatever the incident wavelength spectrum.

# OTHER IMPACTS OF VCN

Specular Reflectometry is 1D technique

Q directions perpendicular to the sample surface are probed



## Optics improvement

- There are no phase space constraints in the direction perpendicular to the incidence plane  
Hence the beam can be focussed as much as possible in this direction
- The efficiency of reflective optics scales as  $m \times 0.1^\circ / \text{\AA}$  where  $m$  is the reflectivity enhancement factor of the super mirrors compared to nickel mirrors.
- If the wavelength spectrum is shifted by a factor  $\alpha$ , the angular phase space which can be used (perpendicular to the incidence plane) is also multiplied by a factor  $\alpha$
- Beware that this has practical limitations
  - With a highly focussing optics in the plane perpendicular to the incidence plane, one has to ensure that all neutrons are collected by the detector.
  - Assuming for example an operation with wavelengths up to  $50 \text{ \AA}$ , and  $m=4$  optics, the critical angle of reflexion is on the order of  $20^\circ$ . If the detector is set at 1.5m from the sample (=focal point), the footprint at the detector position is  $1500 \times 2 \times \sin(20^\circ) = 1000\text{mm}$  !
- Assuming these technical constraints can be overcome, **the gain in flux is proportional to  $\alpha$**

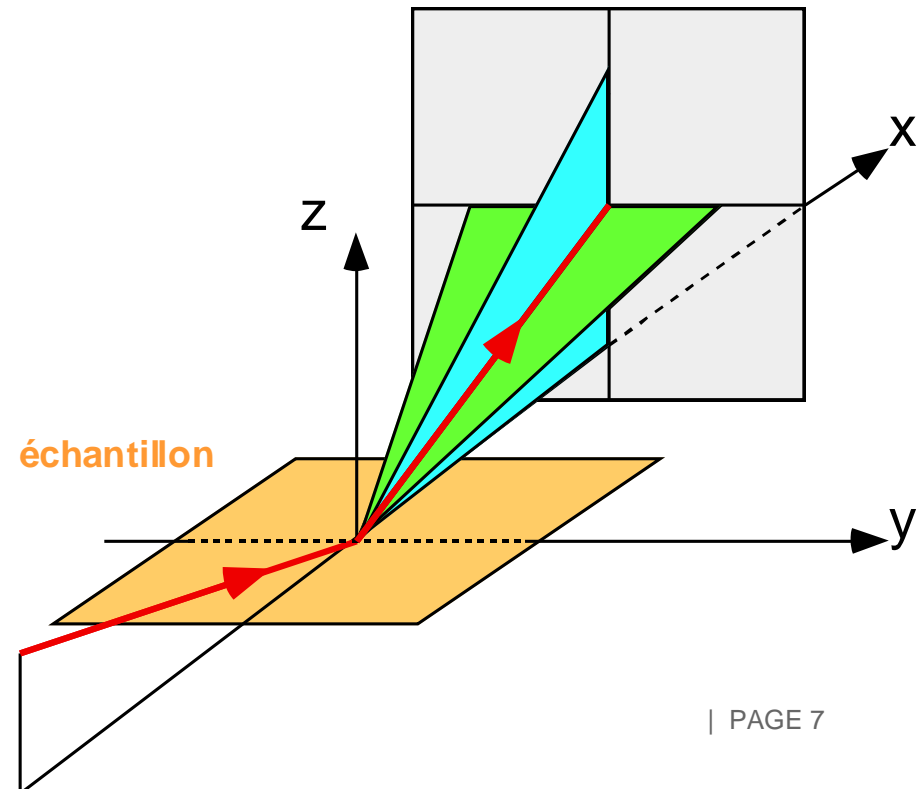
# SPECIFIC CASE OF OFF-SPECULAR AND GISANS

Off-specular scattering

→ the same considerations as for specular reflectometry apply

GISANS

→ the focussing optics cannot be used



# THE ISSUE OF TIME RESOLUTION

The ESS long pulse / low repetition rate is especially suited to low resolution techniques such as reflectometry.

Using VCN introduces extra constraints

Assuming that the wavelength spectrum is shifted by a factor  $\alpha$ ,  
and that the pulse length is kept constant (at 2.6ms),

- The instrument length would be increased by a factor  $\alpha$  to keep the same resolution
- Any increase in the pulse length would lead to a proportionnal increase of the instrument length to compensate
- When the pulse time structure become too sub-optimal, other implementations could be considered but they would be less efficient and reduce the Q-range which can be measured (for a given  $\theta_i$ )
- At the extreme limit, using a continuous VCN source would lead to only marginal improvements compared to ILL instruments.



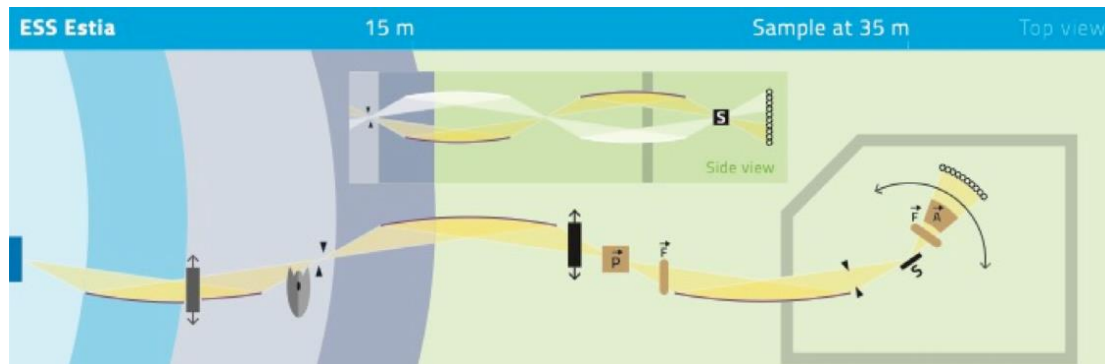
The above considerations apply also for reflectometry implemented on a continuous source

- Either on a fixed wavelength instrument where the operation wavelength can be shifted to higher values
- Or on a chopper TOF instrument for which the very same considerations apply

# OTHER MORE EXOTIC TECHNIQUES

In the previous decades other implementation of reflectometry measurements have been proposed:

- **Energy analysis of a white beam** (via refraction or optics, ex. RAINBOW or EASYREF)  
Such techniques are efficient only on continuous sources but would strongly benefit from a shift to longer wavelengths
  - Optics efficiency is proportional to  $\lambda$  ( $m \times 0.1^\circ / \text{\AA}$ )
  - Refraction efficiency is proportional to  $\lambda^2$
- **SERGIS** techniques using spin-echo techniques
  - Refer to the discussion on the Spin-Echo techniques with VCN
- **SELENE type ( $\rightarrow$  ESTIA@ESS)**
  - Optics efficiency is proportional to the wavelength ( $m \times 0.1^\circ / \text{\AA}$ )
  - Technical implementation is much easier
    - Either shorter optics or broader  $\lambda$  and angular range



The time structure of ESS is ideally suited for reflectometry experiments → hard to compete

Assuming a temperature shift from 20K to 5K,

- The wavelength spectrum would be shifted by a factor 2

Technique	Gain	20K → 5K
Specular reflectivity	$\propto \lambda^2$	4
Off-specular reflectivity	$\propto \lambda^2$	4
GISANS	$\propto \lambda$	2

- Assuming that the integrated VCN flux is identical to the CN flux (brightness value)
- Assuming the time structure is not too dilated (pulse < 10ms)
- VCN neutrons could be practically useful up to a shift of a factor 3 in the wavelength spectrum.
- Beware that neutrons are falling in the gravity field  
→ challenge to achieve clean horizontal instrument measurements